Continuous Ground-Based Observation of Aerosol Optical Properties at Tsukuba, Japan: Trend and Climatology

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Abstract

To investigate aerosol optical properties, the Meteorological Research Institute has been continuously measuring scattering and absorption coefficients since January 2002 by using an integrating nephelometer and one- and three-wavelength absorption photometers in dry air conditions at Tsukuba, Japan. We used these optical data to investigate trends of aerosol properties and climatology from 2002 to 2013. The results showed that most aerosol characteristics had seasonal variation and decreasing or increasing trends significant at the 95% confidence level. From 2002 to 2013, the extinction coefficient at 550 nm and absorption coefficient at 530 nm had statistically significant decreases of $-1.5 \times 10^{-6}$ and $-5.4 \times 10^{-7} \text{m}^{-1} \text{year}^{-1}$, respectively. In the same period, the scattering coefficient showed a non-significant decrease of $-8.8 \times 10^{-7} \text{m}^{-1} \text{year}^{-1}$. The single scattering albedo (SSA) at 550 nm had a significant increasing trend of $7.4 \times 10^{-3} \text{year}^{-1}$. Asymmetry factors did not show a significant trend. The increasing trend in the extinction Ångström exponent was significant, whereas the trend in the effective radius was not significant. The increasing trend of $2.1 \times 10^{-2} \text{year}^{-1}$ in the absorption Ångström exponent from 2006 to 2013 was significant. This tendency suggests a compositional change of light-absorbing aerosol. Frequency distributions of aerosol properties were investigated during 2006–2012. In this period, absorption coefficients were measured by the three-wavelength absorption photometer. The most frequent values of the extinction coefficient at 550 nm, the absorption coefficient at 530 nm, and the SSA at 550 nm were $25 \times 10^{-6}$, $3.0 \times 10^{-6} \text{m}^{-1}$, and 0.905, respectively. The analysis using the extinction Ångström exponent showed that aerosol characteristics were dependent on the extinction Ångström exponent. The aerosol characteristics estimated from optical data were consistent with those derived from radiometer data. Therefore, ground-based monitoring of aerosol optical properties is useful for monitoring aerosol characteristics and interpreting variations in the surface radiation budget.

Keywords aerosol optical property trend; aerosol optical property climatology; single scattering albedo; aerosol scattering coefficient; aerosol absorption coefficient

1. Introduction

Aerosol characteristics are an important factor in the Earth’s radiation budget. The Earth’s radiation budget is influenced by radiatively active gases, aerosols, and clouds. Aerosols change the radiation budget directly by absorbing and scattering solar radiation and indirectly through their role as cloud condensation nuclei (CCN), thereby increasing cloud reflectivity and lifetimes (e.g., Ramanathan et al. 2001; Lohmann and Feichter 2005).
The surface radiation budget has variability at the decadal scale (Wild 2009). Long-term observations of surface solar radiation revealed a decreasing trend, called global dimming, from the early 1950s to the late 1980s (Gilgen et al. 1998; Liepert 2002; Wild et al. 2004; Wild 2009) and an increasing trend, called global brightening, in the 1990s (Wild et al. 2005, 2007, 2008; Pinker et al. 2005; Dutton et al. 2006; Gilgen et al. 2009; Wild 2009). In Japan, likewise, surface solar radiation decreased until the mid-1980s and increased thereafter (Wild et al. 2005; Ohmura 2009), and the rate of increase has decreased more recently (Wild et al. 2009; Kudo et al. 2012). The variation in the observed surface solar radiation depends on the presence of clouds, aerosols, and radiatively active gases.

At the Meteorological Research Institute (MRI), we are investigating the influence of aerosol optical properties through observations of surface solar radiation and direct measurements of aerosol optical properties (scattering coefficients and absorption coefficients) at a ground-based observation site in Tsukuba, Japan. On the basis of surface solar radiation data, Kudo et al. (2011, 2012) reported that decadal variation from the 1970s to 2000s in Japan could be explained by variations in aerosol optical depth and single scattering albedo (SSA) rather than by changes in clouds.

The objectives of this study are to investigate the most recent 10 years of aerosol optical properties measured at Tsukuba and to show the climatological characteristics of the Japan region. These data consist of optical scattering and absorption coefficients of aerosol. The aerosol composition is not measured; therefore, we cannot address the relationship between composition and optical properties. This paper concentrates on the observational evidence bearing on the aerosol optical properties.

In the 1970s, the importance of aerosol optical properties was recognized (e.g., Yamamoto and Tanaka 1972), and measurement programs were begun at several locations (e.g., South Pole, Mauna Loa, Point Barrow; McComiskey et al. 2004). Awareness of the effect of aerosols on climate radiative forcing led to an increase in the number of measured variables and measurement sites in the 1990s.

Coen et al. (2007) analyzed data from the alpine Jungfraujoch station in Switzerland (JFJ; 3580 m asl), where measurements of a comprehensive set of aerosol variables began in 1995, and described an increasing trend in light-scattering coefficients during the September to December period. Junker et al. (2006) reported the trends and seasonal characteristics of aerosol light absorption at the Mace Head Atmospheric Research station near Carna, Ireland, where aerosol measurements based on light attenuation in fiber filters have been conducted since 1989. Sheridan et al. (2001) reported the results of 4 years of continuous measurement of aerosol optical and microphysical properties at the U.S. Department of Energy’s Atmospheric Radiation Measurement (ARM) Program Southern Great Plains Cloud and Radiation Testbed (CART) site.

We began test measurements of aerosol optical properties at the MRI in 2001, and have performed continuous measurements since 2002. Scattering coefficients are measured with a three-wavelength nephelometer (TSI model 3563), and absorption coefficients were initially measured with a Particle Soot/Absorption Photometer (PSAP), and beginning in 2006, they have been measured with a three-wavelength PSAP (PSAP3λ) (both by Radiance Research).

In this paper, the data we used are described in Section 2, and analytical methods are described in Section 3. Section 4 presents the 10-year trends in aerosol optical properties, frequency distribution of aerosol optical properties, and an analysis of aerosol properties using the extinction Ångström exponent. Our results are summarized in Section 5.

2. Data

Scattering coefficients were measured from January 2002 to July 2013 by a nephelometer (TSI model 3563, described by Anderson et al. 1996) using 30-min averages. The nephelometer measures scattering coefficients and hemispheric back-scattering coefficients at wavelengths of 450, 550, and 700 nm. The measurement angle ranges of scattering coefficients and hemispheric back-scattering coefficients are 7° to 170° and 90° to 170°, respectively. Because aerosol scattering coefficients are affected by relative humidity, it is difficult to measure them reliably in open air. Therefore, we measured scattering coefficients in dry air conditions when relative humidity was less than 50%.

Absorption coefficients were determined from PSAP and PSAP3λ measurements using 30-min averages. The measurement periods of PSAP and PSAP3λ were from January 2002 to July 2013 and from June 2006 to July 2013, respectively. The measurement wavelength was 565 nm for PSAP and 467, 530, and 660 nm for PSAP3λ. The values measured by PSAP and PSAP3λ were corrected using Bond et al.’s (1999) method.

The integrating nephelometer is widely used for
measuring aerosol scattering coefficients; however, it cannot measure light scattered in the extreme forward or backward directions (scattering angles near 0° and 180°; Heintzenberg and Charlson 1996; Anderson et al. 1996; Anderson and Ogren 1998; Müller et al. 2009). To correct this truncation error, information on aerosol absorption properties and the particle size distribution is necessary (Bond et al. 2009). We used a recently developed method that uses multiwavelength scattering and absorption coefficient data to calculate corrected scattering coefficients (see Section 3) (Uchiyama 2014) and obtained more accurate values for scattering coefficients and other single scattering properties (extinction coefficients, SSA, and asymmetry factor). To analyze single-wavelength PSAP data, together with three-wavelength PSAP3λ data, absorption coefficients were extrapolated by assuming wavelength dependence $\lambda^{-\alpha}$. In this study, $\alpha = 1.02 (\pm 0.16)$, the average value of monthly average (standard deviation) from March 2006 to July 2013, was used.

The continuous measurements were performed at observation building for surface radiation measurements at MRI (36.06°N, 140.01°E) in Tsukuba, Japan. Tsukuba is approximately 60 km northeast of the center of Tokyo and has a population of about 200,000 with a population density of about 750 people km$^{-2}$. Regional land is used for vegetable and rice fields, except in the central city area. There is no large source of anthropogenic emissions near the observation site.

Sample air was drawn into the building through an intake 3 m above the ground by using a stainless steel tube with an inner diameter of 20 mm. Room temperature was controlled by weak air conditioning. Therefore, it was slightly cooler (higher) than the outside air temperature in summer (winter), and was almost the same as the outside air temperature in spring and autumn. The air was heated to 50°C–60°C near the entrance of the intake. Inside the building, the air was divided and introduced into each instrument through conductive silicone tubes covered with heat insulating material. These steps ensured that the sample air met the necessary condition of low relative humidity.

3. Methods

3.1 Data analysis

The nephelometer data require corrections for the angular truncation error and the light source distribution error. Our correction used a statistical method to simultaneously retrieve the complex refraction index (both real and imaginary parts) and the aerosol size distribution from the three-wavelength scattering and hemispheric back-scattering coefficients and three-wavelength absorption coefficients (Uchiyama 2014). In this method, we seek the solution that minimizes the following function:

$$J(x) = \frac{1}{2}(x - x_0)'C_0^{-1}(x - x_0) + \frac{1}{2}(f(x) - f_{\text{obs}})'C_{\text{obs}}^{-1}(f(x) - f_{\text{obs}}),$$

where vector $x$ is a parameter to be retrieved, vector $x_0$ is the first estimate of the parameter, vector $f_{\text{obs}}$ is the measured data, $f(x)$ is the measured value of $x$, and $x'$ is the transpose of vector $x$. It is assumed that the errors of vectors $x_0$ and $f_{\text{obs}}$ are uncorrelated and follow the normal distribution of the covariance matrices $C_0$ and $C_{\text{obs}}$.

Once the complex refraction index and aerosol size distribution are retrieved, single scattering properties (extinction, scattering and absorption coefficients, SSA, phase function, and asymmetry factor) can be calculated. The resulting scattering coefficients are the corrected values. Results of a numerical simulation showed that the root mean square errors (RMSEs) of the SSA calculated directly from the measured values are 0.014–0.021 and that of the SSA calculated from the corrected values is 0.002, corresponding to a relative error of 0.2% (Uchiyama 2014).

3.2 Calculation of trends

Some aerosol properties displayed seasonal variation as well as long-term trends. We removed the seasonal variation by assuming the simple form of the following equation:

$$f(t) = (a \cdot t + b) \cdot (c_0 + c_1 \sin(\omega t) + c_2 \cos(\omega t) + c_3 \sin(2\omega t) + c_4 \cos(2\omega t)),$$

where $\omega = 2\pi / T$ and $T=1$ year. In this equation, periods of 1 year and 0.5 year are taken into account, and the long-term trend is $a \cdot c_0$. The coefficients are determined by the following procedure. First, coefficients $a$ and $b$ are determined by the least squares method using all data. Second, coefficients $c_0$, $c_1$, $c_2$, $c_3$, and $c_4$ are determined from $f(t) / (a \cdot t + b)$ by the least squares method. This procedure is repeated several times to reduce the residual errors. Our analysis of the long-term trend used monthly average values.
4. Results

4.1 Trends of aerosol properties

The trends of aerosol properties are summarized in Table 1. The time series of scattering coefficients at a wavelength of 550 nm and absorption coefficients at 530 nm are shown in Fig. 1. The absorption coefficients were fitted well by Eq. (2) in contrast to the scattering coefficients, which did not (figures not shown). The absorption coefficients showed a clear seasonal variation, but the scattering coefficients did not. The absorption coefficients were large in the late autumn and early winter and small in the summer. The trends of scattering coefficients from 2002 to 2013 were $-9.7 \times 10^{-7}$, $-8.8 \times 10^{-7}$, and $-6.6 \times 10^{-7}$ m$^{-1}$ year$^{-1}$ at wavelengths of 450, 550, and 700 nm, respectively. The trend of absorption coefficients from 2002 to 2013 was $-5.4 \times 10^{-7}$ m$^{-1}$ year$^{-1}$ at 530 nm, and those from 2006 to 2013 were $-6.4 \times 10^{-7}$, $-5.6 \times 10^{-7}$, and $-4.7 \times 10^{-7}$ m$^{-1}$ year$^{-1}$ at 467, 530, and 660 nm, respectively. Both scattering and absorption coefficients showed a decreasing trend, but only the trend of the absorption coefficients was significant at the 95 % confidence level. A similar decrease in the absorption coefficients was significant at the 95 % confidence level. A similar decrease in the absorption coefficients was significant at the 95 % confidence level.

Table 1. Trends of aerosol properties.

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<td>Cabs(467nm)</td>
<td>$(7.6 \times 10^{-7})$</td>
<td>$(-9.9 \times 10^{-7})$</td>
<td>$(-6.1 \times 10^{-7})$</td>
<td>$(-6.4 \times 10^{-7})$</td>
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<tr>
<td>Cabs(530nm)</td>
<td>$6.7 \times 10^{-7}$</td>
<td>$-8.7 \times 10^{-7}$</td>
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<tr>
<td>Cabs(660nm)</td>
<td>$5.3 \times 10^{-7}$</td>
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<td>$-4.7 \times 10^{-7}$</td>
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<tr>
<td>Cext(450nm)</td>
<td>$-6.5 \times 10^{-6}$</td>
<td>$-2.8 \times 10^{-6}$</td>
<td>$-1.7 \times 10^{-6}$</td>
<td>$-2.5 \times 10^{-6}$*</td>
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<tr>
<td>Cext(550nm)</td>
<td>$-4.8 \times 10^{-6}$</td>
<td>$-2.2 \times 10^{-6}$</td>
<td>$-1.5 \times 10^{-6}$</td>
<td>$-2.0 \times 10^{-6}$*</td>
</tr>
<tr>
<td>Cext(700nm)</td>
<td>$-3.2 \times 10^{-6}$</td>
<td>$-1.4 \times 10^{-6}$</td>
<td>$-1.1 \times 10^{-6}$</td>
<td>$-1.2 \times 10^{-6}$*</td>
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<td>Csca(450nm)</td>
<td>$-7.3 \times 10^{-6}$</td>
<td>$-1.8 \times 10^{-6}$</td>
<td>$-9.7 \times 10^{-7}$</td>
<td>$-1.9 \times 10^{-6}$*</td>
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<tr>
<td>Csca(550nm)</td>
<td>$-5.5 \times 10^{-6}$</td>
<td>$-1.4 \times 10^{-6}$</td>
<td>$-8.8 \times 10^{-7}$</td>
<td>$-1.4 \times 10^{-6}$*</td>
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<td>Csca(700nm)</td>
<td>$-3.8 \times 10^{-6}$</td>
<td>$-7.7 \times 10^{-7}$</td>
<td>$-6.6 \times 10^{-7}$</td>
<td>$-7.7 \times 10^{-7}$*</td>
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<tr>
<td>SSA(450nm)</td>
<td>$-1.6 \times 10^{-2}$</td>
<td>$8.6 \times 10^{-3}$</td>
<td>$7.2 \times 10^{-3}$</td>
<td>$6.0 \times 10^{-3}$</td>
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<tr>
<td>SSA(550nm)</td>
<td>$-1.7 \times 10^{-2}$</td>
<td>$9.3 \times 10^{-3}$</td>
<td>$7.4 \times 10^{-3}$</td>
<td>$6.9 \times 10^{-3}$</td>
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<tr>
<td>SSA(700nm)</td>
<td>$-2.0 \times 10^{-2}$</td>
<td>$1.1 \times 10^{-2}$</td>
<td>$7.9 \times 10^{-3}$</td>
<td>$8.8 \times 10^{-3}$</td>
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<tr>
<td>Asymmetry factor(450nm)</td>
<td>$1.4 \times 10^{-3}$</td>
<td>$2.2 \times 10^{-3}$</td>
<td>$0.4 \times 10^{-3}$</td>
<td>$3.5 \times 10^{-3}$</td>
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<td>Asymmetry factor(550nm)</td>
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<td>$2.4 \times 10^{-3}$</td>
<td>$0.2 \times 10^{-3}$</td>
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<tr>
<td>Asymmetry factor(700nm)</td>
<td>$2.6 \times 10^{-3}$</td>
<td>$1.1 \times 10^{-3}$</td>
<td>$-1.3 \times 10^{-3}$*</td>
<td>$2.5 \times 10^{-3}$**</td>
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<tr>
<td>Alpha_ext</td>
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<td>$-4.5 \times 10^{-3}$</td>
</tr>
<tr>
<td>Alpha_sca</td>
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<td>$-3.5 \times 10^{-3}$</td>
<td>$1.1 \times 10^{-2}$*</td>
<td>$-1.3 \times 10^{-2}$</td>
</tr>
<tr>
<td>Reff</td>
<td>$-1.2 \times 10^{-3}$</td>
<td>$1.5 \times 10^{-3}$</td>
<td>$-0.5 \times 10^{-3}$</td>
<td>$3.6 \times 10^{-3}$*</td>
</tr>
<tr>
<td>Veff</td>
<td>$1.2 \times 10^{-1}$</td>
<td>$-4.0 \times 10^{-2}$</td>
<td>$-1.0 \times 10^{-1}$*</td>
<td>$6.1 \times 10^{-2}$</td>
</tr>
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</table>

Asterisk (*) means that the trend is significant at the 95 % confidence level.
Two asterisks (**) mean that the trend is significant at the 90 % confidence level.
1WL means that one-wavelength PSAP was used as the data. 3WL means that three-wavelength PSAP3$\lambda$ was used.
Cabs: absorption coefficient (m$^{-1}$), Alpha_abs: Ångström exponent for Cabs, Cext: extinction coefficient (m$^{-1}$), Csca: scattering coefficient (m$^{-1}$), SSA: single scattering albedo, Alpha_ext: Ångström exponent for Cext, Alpha_sca: Ångström exponent for Csca, Reff: effective radius (µm), Veff: effective variance.
confidence level during 2002–2013 and significant at the 90% confidence level during 2006–2013.

The time series of the SSA at 550 nm is shown in Fig. 2. Most SSA values lie between 0.8 and 0.9. In a clear pattern of seasonal variation, the SSA was large (relative absorption is less) in spring and early summer and small in late autumn and early winter. Trends of the SSA from 2002 to 2013 were $7.2 \times 10^{-3}$, $7.4 \times 10^{-3}$, and $7.9 \times 10^{-3}$ year$^{-1}$ at 450, 550, and 700 nm, respectively, and those from 2006 to 2013 were $6.0 \times 10^{-3}$, $6.9 \times 10^{-3}$, and $8.8 \times 10^{-3}$ year$^{-1}$. These trends were significant at the 95% confidence level. The seasonal variation in the SSA was also apparent in the radiometer data (Nishizawa et al. 2004; Kudo et al. 2010), as was the increase in the SSA (Kudo et al. 2010, 2012).

If the decreasing trends of extinction coefficients and the increasing trends of the SSA occur in the whole atmosphere, these trends cause an increase in the downward solar irradiance at the Earth’s surface (Kudo et al. 2010).

Asymmetry factor is defined as follows (Liou 2002):
where $P(\cos \Theta)$ is the phase function of scattering, and $\Theta$ is the scattering angle. Asymmetry factor varies between 1 and $-1$, is 0 for isotropic scattering, and can usefully characterize the shape of the phase function of scattering.

The time series of asymmetry factor at 550 nm is shown in Fig. 3. The values of asymmetry factor varied between 0.5 and 0.65 and displayed a clear seasonal variation, being large in summer and small in late autumn and early winter. The trends of asymmetry factor from 2002 to 2013 were $0.4 \times 10^{-3}$, $0.2 \times 10^{-3}$, and $-1.3 \times 10^{-3}$ year$^{-1}$ at 450, 550, and 700 nm, respectively, and those from 2006 to 2013 were $3.5 \times 10^{-3}$, $3.7 \times 10^{-3}$, and $2.5 \times 10^{-3}$ year$^{-1}$. The trends from 2002 to 2013 were not significant at the 95 % confidence level, whereas those from 2006 to 2013 were significant at the 90 % or 95 % confidence level, depending on wavelength. Asymmetry factors from 2008 to 2010 were a little smaller than those in the other periods, which may affect the increasing trends for 2006–2013. A lower asymmetry factor means that the relative contribution of smaller particles to the size distribution is greater. Because the measurement procedure did not change between 2006 and 2013, this change in the particle size distribution can be regarded as real. The trends at 700 nm for both 2002–2006 and 2006–2013 were positive (increasing), but the trend at 700 nm for 2002–2013 was negative (decreasing). Though the time series of asymmetry factor at 700 nm is not shown here, this decreasing trend was caused by the lower asymmetry factor from 2011 to 2013 than from 2002 to 2006. This tendency can be seen in Fig. 3.

The wavelength dependence of the extinction coefficients is approximated as follows:

$$C_{ext} \propto \lambda^{-\alpha}. \tag{4}$$

The exponent $\alpha$ called the extinction Ångström exponent, is an index of the size distribution; when $\alpha$ is large, the size distribution favors small particles, and when $\alpha$ is small, the size distribution favors large particles.

The time series of extinction Ångström exponents is shown in Fig. 4. The value varied mostly between 1.3 and 1.8 and did not show seasonal variation. When airborne mineral dust is abundant over the observation site, the extinction Ångström exponent becomes less than 0.5 for a day or so. However, monthly average values are not affected by dust events due to the short duration of dust event. The trend of extinction Ångström exponent was $1.4 \times 10^{-2}$ year$^{-1}$ from 2002 to 2013 and $-4.5 \times 10^{-3}$ year$^{-1}$ from 2006 to 2013. The trend from 2002 to 2013 was significant at the 95 % confidence level. After the summer of 2012, periods of lower extinction Ångström exponent were much more common. This accounts for the decreasing tendency from 2006 to 2013. Usually, an increase in the extinction Ångström exponent means an increase in smaller particles, which should correspond to a decrease in the asymmetry factor. The trends in asymmetry factor and extinction Ångström exponent were consistent from 2006 to 2013, but for 2002 to 2013, this relationship was not clear.

We can analogously define an Ångström exponent for absorption coefficients to the extinction Ångström exponent as follows:

$$C_{abs} \propto \lambda^{-\alpha}. \tag{5}$$

The time series of absorption Ångström exponents is shown in Fig. 5. The value varied mostly between 0.8 and 1.2 and displayed clear seasonal variation. The trend of the absorption Ångström exponent from 2006 to 2012 was $2.1 \times 10^{-2}$ year$^{-1}$ and was significant at the 95 % confidence level. The values in 2008 and 2009 ranged from 0.8 to 1 and were lower than those in other years.

The absorption Ångström exponent depends on the source of emission, aerosol composition, and the stage of aerosol aging (Kirchstetter et al. 2004; Bergstrom et al. 2007; Lewis et al. 2008; Yang et al. 2009; Russell et al. 2010). Light absorption by motor vehicle aerosols exhibits a relatively weak wavelength dependence; the absorption Ångström exponent is approximately 1, indicating that black carbon is the dominant absorbing aerosol component. Biomass smoke aerosols from biomass burning have much stronger wavelength dependence, with absorption Ångström exponents of approximately 1.5 to 2.5, as a result of enhanced light absorption at wavelengths shorter than 600 nm due to the organic carbon component. Mineral dust aerosol also has stronger wavelength dependence, with absorption Ångström exponents of approximately 1.5 to 3.0.

In our data, the absorption Ångström exponent tended to increase. As lower absorption Ångström exponents are attributable to motor vehicles, the restriction of diesel engines in the Tokyo area since 2003 may have caused a decrease in aerosols with lower absorption Ångström exponents; thus, the
relative contribution of aerosols with higher absorption Ångström exponents may have increased. The average value of the absorption Ångström exponent from 2006 to 2012 was about 1.02. The values in 2008 and 2009 were 0.8 to 1. It is possible that a local source near the observation site produced aerosols with characteristics different from those of the typical aerosol during these years.

Figure 5 shows a clear seasonal variation in the absorption Ångström exponent, which was small in summer (about 1) and large in late autumn and early winter (about 1.15). This may reflect seasonal local biomass burning in the late autumn and early winter.

The effective radius and variance are defined as follows (Hansen and Travis 1974):

$\text{r}_{\text{eff}} = \frac{\int \pi r^2 n(r) \, dr}{\int \pi r^2 n(r) \, dr}$

$\text{v}_{\text{eff}} = \frac{\int (r - \text{r}_{\text{eff}})^2 \pi r^2 n(r) \, dr}{\text{r}_{\text{eff}}^2 \int \pi r^2 n(r) \, dr}$

The time series of effective radius ($\text{r}_{\text{eff}}$) is shown in Fig. 6. The values of monthly average $\text{r}_{\text{eff}}$ varied between 0.08 and 0.2, tending to be larger (with larger standard deviations) in summer and smaller in winter (with smaller standard deviations), except for the summers of 2008 and 2009. The trend of $\text{r}_{\text{eff}}$ from 2002 to 2013 was $-0.5 \times 10^{-3} \, \mu m \, year^{-1}$ (not significant at the 95 % confidence level) and that from 2006 to 2013 was $3.6 \times 10^{-3} \, \mu m \, year^{-1}$ (significant at the 95 % confidence level). After the summer of 2012, periods of larger effective radius were much more common. This accounts for the increasing tendency from 2006 to 2013. Excluding these period data, the tendency was negative (decreasing). The tendencies of asymmetry factor, extinction Ångström exponent and effective radius for 2006 to 2013 were consistent; these tendencies correspond to an increase in larger particles. However, the tendencies of these parameters for 2002 to 2013 were not clear.
The time series of effective variance (veff) is shown in Fig. 7. The values of monthly average veff varied between 2 and 6, with a large scatter and standard deviations of about 2. The trend of veff from 2002 to 2013 was \(-1.0 \times 10^{-1}\) year\(^{-1}\) and that from 2006 to 2013 was \(6.1 \times 10^{-2}\) year\(^{-1}\). The 2002–2013 trend was significant at the 95 % confidence level.

4.2 Climatology

This section presents the frequency distribution of aerosol optical properties (extinction coefficient at 550 nm, absorption coefficients at 530 nm, SSA at 550 nm, and asymmetry factor at 550 nm) based on the data for 2006–2013. For the period 2006–2013, we used three-wavelength PSAP3\(\lambda\) data. This section also presents optical properties classified by the magnitude of extinction Ångström exponent.

a. Frequency distribution

The frequency distribution of extinction coefficients using all data is shown in Fig. 8a, and a seasonal breakdown is shown in Fig. 8b. The most frequent value was \(25 \times 10^{-6}\) m\(^{-1}\), and most of the data were below \(200 \times 10^{-6}\) m\(^{-1}\). The most frequent values in spring (March to May) and summer (June to August) were largest at \(35 \times 10^{-6}\) m\(^{-1}\), the value in autumn (September to November) was \(25 \times 10^{-6}\) m\(^{-1}\), and that in winter (December to February) was smallest at \(15 \times 10^{-6}\) m\(^{-1}\). In winter, the smaller extinction coefficients were observed when cold air masses flowed from Siberia to Japan. Values greater than \(150 \times 10^{-6}\) m\(^{-1}\) were more frequent in winter and autumn than in spring and summer. Night-time temperature inversions in winter and late autumn tend to accumulate aerosols and increase extinction coefficients. The monthly average values of some aerosol properties are shown in Table 2. The monthly average values of the extinction and absorption coefficients in Table 2 show the same characteristics as those shown in Fig. 8.

The frequency distribution of absorption coefficients is shown in Fig. 9. The most frequent value in all seasons was \(3.0 \times 10^{-6}\) m\(^{-1}\), and most of the data were below \(30 \times 10^{-6}\) m\(^{-1}\). The frequency of absorption coefficients greater than \(15 \times 10^{-6}\) m\(^{-1}\) was higher in autumn and winter than in spring and summer. Inversion layers have the same effect on absorption coefficients as they do on extinction coefficients. The most frequent value does not depend on the season. This suggests that most of the absorbing aerosols are emitted from the area near the observation site.

The frequency distribution of the SSA is shown in Fig. 10. The most frequent value in the entire data was near 0.905. The most frequent SSA values in spring, summer, autumn, and winter were 0.925, 0.925, 0.885, and 0.875, respectively. The seasonal effect in
Fig. 8. Frequency distribution of extinction coefficient. Data in the period from 2006 to 2013 were used. (a) Frequency distribution of extinction coefficient for entire data. (b) Frequency distribution of extinction coefficient for each season: spring (March to May), summer (June to August), autumn (September to November), and winter (December to February).

Table 2. Monthly averages of aerosol properties.

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<th>Month</th>
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<td>SG</td>
<td>Cabs (530 nm) ×10⁻⁶</td>
<td>SG</td>
<td>Cabs (530 nm) ×10⁻⁶</td>
<td>SG</td>
<td>Cabs (530 nm) ×10⁻⁶</td>
<td>SG</td>
<td>Cabs (530 nm) ×10⁻⁶</td>
<td>SG</td>
<td>Cabs (530 nm) ×10⁻⁶</td>
<td>SG</td>
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<tr>
<td>1</td>
<td>8.34</td>
<td>(7.52)</td>
<td>8.43</td>
<td>(6.69)</td>
<td>6.00</td>
<td>(4.05)</td>
<td>5.84</td>
<td>(3.62)</td>
<td>5.69</td>
<td>(4.30)</td>
<td>5.47</td>
<td>(2.85)</td>
</tr>
<tr>
<td>Alpha_abs</td>
<td>1.07</td>
<td>(0.24)</td>
<td>1.07</td>
<td>(0.20)</td>
<td>1.08</td>
<td>(0.17)</td>
<td>1.08</td>
<td>(0.18)</td>
<td>0.97</td>
<td>(0.24)</td>
<td>1.01</td>
<td>(0.25)</td>
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<tr>
<td>Cext (550 nm) ×10⁻⁶</td>
<td>51.99</td>
<td>(40.83)</td>
<td>71.47</td>
<td>(53.95)</td>
<td>65.81</td>
<td>(37.52)</td>
<td>62.51</td>
<td>(36.71)</td>
<td>62.51</td>
<td>(37.52)</td>
<td>48.03</td>
<td>(31.23)</td>
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<tr>
<td>Csca (550 nm) ×10⁻⁶</td>
<td>43.77</td>
<td>(34.84)</td>
<td>62.91</td>
<td>(48.39)</td>
<td>59.62</td>
<td>(40.06)</td>
<td>56.85</td>
<td>(34.44)</td>
<td>63.62</td>
<td>(38.71)</td>
<td>63.62</td>
<td>(38.71)</td>
</tr>
<tr>
<td>SSA (550 nm)</td>
<td>0.844</td>
<td>(0.067)</td>
<td>0.877</td>
<td>(0.053)</td>
<td>0.901</td>
<td>(0.041)</td>
<td>0.901</td>
<td>(0.042)</td>
<td>0.906</td>
<td>(0.045)</td>
<td>0.906</td>
<td>(0.045)</td>
</tr>
<tr>
<td>Asym.factor (550 nm)</td>
<td>0.564</td>
<td>(0.036)</td>
<td>0.581</td>
<td>(0.034)</td>
<td>0.593</td>
<td>(0.032)</td>
<td>0.596</td>
<td>(0.030)</td>
<td>0.598</td>
<td>(0.034)</td>
<td>0.607</td>
<td>(0.032)</td>
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<tr>
<td>Alpha_ext</td>
<td>1.66</td>
<td>(0.21)</td>
<td>1.62</td>
<td>(0.24)</td>
<td>1.59</td>
<td>(0.26)</td>
<td>1.53</td>
<td>(0.29)</td>
<td>1.60</td>
<td>(0.27)</td>
<td>1.59</td>
<td>(0.27)</td>
</tr>
<tr>
<td>Reff</td>
<td>0.093</td>
<td>(0.015)</td>
<td>0.107</td>
<td>(0.027)</td>
<td>0.118</td>
<td>(0.033)</td>
<td>0.130</td>
<td>(0.047)</td>
<td>0.118</td>
<td>(0.034)</td>
<td>0.114</td>
<td>(0.049)</td>
</tr>
<tr>
<td>Veff</td>
<td>3.52</td>
<td>(1.67)</td>
<td>4.13</td>
<td>(2.17)</td>
<td>4.50</td>
<td>(2.25)</td>
<td>4.82</td>
<td>(2.41)</td>
<td>4.29</td>
<td>(2.25)</td>
<td>4.61</td>
<td>(2.25)</td>
</tr>
</tbody>
</table>

Cabs (530 nm): absorption coefficient at 530 nm (m⁻¹), SD: standard deviation, Alpha_abs: Ångström exponent for Cabs, Cext (550 nm): extinction coefficient at 550 nm (m⁻¹), Csca (550 nm): scattering coefficient at 550 nm (m⁻¹), SSA (550 nm): single scattering albedo at 550 nm, Asym.factor (550 nm): Asymmetry factor at 550 nm, Alpha_ext: Ångström exponent for Cext, Reff: effective radius (µm), Veff: effective variance.
late autumn and early winter reflects the burning of field biomass and fallen leaves near the observation site and the increased combustion of heating oil and gas in winter. The monthly average values of the SSA in Table 2 show the same features.

The frequency distribution of the asymmetry factor is shown in Fig. 11. The most frequent value in the entire data was 0.585. The asymmetry factor was smaller in winter and autumn, highest in summer, and intermediate in spring. The marine air mass prevalent in summer increases the proportion of larger particles such as sea salt particles. Furthermore, the relative humidity in summer is higher than that in the other seasons, such that hygroscopic growth affects particle sizes. The sample air in spring includes more mineral dust from deserts in continental Asia than in autumn and winter. Therefore, the asymmetry factor is generally larger in spring than in autumn and winter.

From the above results on frequency distribution of aerosol properties, the variations in aerosol properties measured on the ground-based observation site were determined by the difference of synoptic scale
air mass, absorbing aerosol emitted from near local source and the vertical structure of the atmosphere.

b. Optical characteristics classified by extinction Ångström exponent

The extinction Ångström exponent represents the relative contribution of larger and smaller size particles to the size distribution, being small when the relative contribution of larger particles is great and large when the relative contribution of smaller particles is great. To investigate the dependence of aerosol characteristics on size distribution, we classified the data by the magnitude of the extinction Ångström exponent as follows: $-0.5 < \alpha_{\text{ext}} \leq 0.5$, $0.5 < \alpha_{\text{ext}} \leq 1.0$, $1.0 < \alpha_{\text{ext}} \leq 1.5$, $1.5 < \alpha_{\text{ext}} \leq 2.0$, and $2.0 < \alpha_{\text{ext}} \leq 2.5$.

Figure 12a shows the SSA plotted against the

![Figure 11](image1.png)

Fig. 11. Same as Fig. 8, except for asymmetry factor at 550 nm.

![Figure 12](image2.png)

Fig. 12. (a) Relation between extinction Ångström exponent and SSA. (b) Relation between extinction Ångström exponent and asymmetry factor. Blue line indicates 450 nm, green line indicates 550 nm, and red line indicates 700 nm.
extinction Ångström exponent. As the extinction Ångström exponent increases, the SSA decreases, and this tendency is dependent on wavelength: at smaller extinction Ångström exponents, the SSA decreases (absorption increases) at shorter wavelengths, and at larger extinction Ångström exponents, the SSA decreases at longer wavelengths. At times when airborne mineral dust is abundant at the observation site, extinction Ångström exponents are less than 0.5 (Uchiyama et al. 2005). Therefore, this wavelength dependence of the SSA appears to be a response to mineral dust levels. In Tsukuba, a large fraction of extinction Ångström exponents was more than 1 and that of absorption Ångström exponents was about 1. Therefore, the relative contribution of absorption to extinction coefficient increases at longer wavelengths.

Figure 12b shows the relation between the extinction Ångström exponent and the asymmetry factor. When the extinction Ångström exponent is small, the relative contribution of larger particles to the size distribution is greater, and vice versa. Therefore, as the extinction Ångström exponent increases, the asymmetry factor decreases. Furthermore, the wavelength dependence becomes weaker at smaller extinction Ångström exponents.

Figure 13 shows the derived volume size distributions (dV/d log r). For extinction Ångström exponents greater than 1, the volume size distributions are unimodal with a peak at around 0.1 µm radius. For extinction Ångström exponents less than 1, the retrieved volume size distributions are bimodal with peaks at around 0.1 and 2 µm. Furthermore, as the extinction Ångström exponent decreases, the fraction of larger particles in the size distribution increases. A similar bimodal volume size distribution has been documented from sky radiometer data at times of heavy airborne mineral dust (Tanaka et al. 1989; Shiobara et al. 1991; Uchiyama et al. 2005).

The estimated index of refraction is shown in Fig. 14. The real part of the refractive index is around 1.5, and the wavelength dependence is smaller in each case. The wavelength dependence of the imaginary part of the refractive index is dependent on the extinction Ångström exponent. For extinction Ångström exponents less than 1, the imaginary part decreases as wavelength increases, whereas at extinction Ångström exponents greater than 1.5, the imaginary part increases as wavelength increases. The imaginary part is less at smaller extinction Ångström exponents than at larger extinction Ångström exponents. The magnitude and wavelength dependence of the imaginary part of the refractive index at smaller extinction Ångström exponents corresponds to the characteristics of mineral dust (Aoki et al. 2005).

5. Conclusions

Aerosol scattering and absorption coefficients have been continuously measured since January 2002 at Tsukuba, Japan, using an integrating nephelometer and absorption photometers. We analyzed this 10-year data record to extract aerosol characteristics.

We investigated aerosol characteristics including scattering and absorption coefficients, the SSA, the asymmetry factor, the extinction Ångström exponent, and the absorption Ångström exponent. Many of these characteristics had seasonal variations as well as long-term trends that were significant at the 95% confidence level.

Both extinction and absorption coefficients had statistically significant decreases (95% confidence level) between 2002 and 2013: the trend for extinction coefficient at 550 nm was $-1.5 \times 10^{-6}$ m$^{-1}$ year$^{-1}$ and that of absorption coefficient at 530 nm was $5.4 \times 10^{-7}$ m$^{-1}$ year$^{-1}$. Scattering coefficients at 550 nm also decreased over this period, but the level of statistical significance was lower: the trend was $8.8 \times 10^{-7}$ m$^{-1}$ year$^{-1}$. These decreasing trends are similar to those previously obtained from analysis of radiometer data.

The SSA at 550 nm increased by $7.4 \times 10^{-3}$ year$^{-1}$ from 2002 to 2013 and by $6.9 \times 10^{-3}$ year$^{-1}$ from 2006 to 2013. These trends were significant at the 95% confidence level. The SSA was high in spring and early summer and low in late autumn and early winter. These trends and seasonal variations are similar to those previously obtained from analysis of radiometer data.

The asymmetry factor trends were not significant at the 95% confidence level. The extinction Ångström exponent increased by $1.4 \times 10^{-2}$ year$^{-1}$, which was significant at the 95% confidence level.

The absorption Ångström exponent values were about 1.0 and showed seasonal variation, with larger values in late autumn and early winter and smaller values in summer. The increase in late autumn and early winter is explained by local biomass burning. The absorption Ångström exponent increased from 2006 to 2013 by $2.1 \times 10^{-2}$ year$^{-1}$, probably because of reductions in motor vehicle aerosols with low absorption Ångström exponents.

Using data from 2006 to 2013, the frequency distributions of aerosol properties were investigated. In this period, absorption coefficients were measured by three-wavelength PSAP3λ.

The most frequent value of extinction coefficients
Fig. 13. Relation between extinction Ångström exponent and volume size distribution.

Fig. 14. Relation between extinction Ångström exponent and refractive index. (a) real part of refractive index, (b) imaginary part of refractive index.
was $25 \times 10^{-6}$ m$^{-1}$, and most values were below $200 \times 10^{-6}$ m$^{-1}$. The most frequent value of absorption coefficients was $3.0 \times 10^{-6}$ m$^{-1}$, and most values were below $30 \times 10^{-6}$ m$^{-1}$. The most frequent value in each season was not changed. The most frequent value of the SSA at 550 nm was 0.905. There was a seasonal pattern, with the most frequent SSA values in spring, summer, autumn, and winter being 0.925, 0.925, 0.885, and 0.875, respectively. The most frequent value of asymmetry factor at 550 nm was 0.585.

We also analyzed the dependence of aerosol characteristics on size distribution, as represented by the extinction Ångström exponent. The SSA tended to decrease as the extinction Ångström exponent increased, and this tendency depended on wavelength. These characteristics reflect the composition of the aerosol. At extinction Ångström exponents greater than 1, the volume size distributions were unimodal with a peak at around 0.1 µm radius, and at extinction Ångström exponents less than 1, the volume size distributions were bimodal with peaks at around 0.1 and 2 µm. The real part of the refractive index was around 1.5, and the wavelength dependence was smaller in each case. The wavelength dependence of the imaginary part of the refractive index was dependent on the extinction Ångström exponent. At extinction Ångström exponents greater than 1, the imaginary part decreased as wavelength increased, and at extinction Ångström exponents greater than 1.5, the imaginary part increased as wavelength increased. The values of the imaginary part were smaller in the former case than in the latter case.

Aerosol characteristics estimated from multi-wavelength integrating nephelometer and absorption photometer data in this study are consistent with those derived from radiometer data. Therefore, ground-based monitoring of aerosol optical properties is useful for monitoring aerosol characteristics and interpreting variations in the surface radiation budget. As optical properties depend on the composition and shape of aerosols, it is necessary to monitor these aspects of aerosols as well as their scattering and absorption coefficients.

**Acknowledgment**

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